

# First structure equation

- Spin connection

Let us consider the differential of the vielbein

$de^a = \frac{1}{2}(\partial_\mu e_\nu^a - \partial_\nu e_\mu^a) dx^\mu \wedge dx^\nu$  it is not a Lorentz vector. Introduce the spin connection connection one form

$$\omega'^a{}_b = \Lambda^{-1 a}{}_c d\Lambda^c{}_b + \Lambda^{-1 a}{}_c \omega^c{}_d \Lambda^d{}_b$$

The quantity  $de^a + \omega^a{}_b \wedge e^b \equiv T^a$ .

transforms as a vector  $T'^a = \Lambda^{-1 a}{}_b T^b$ .

torsion tensor  $T^a_{\mu\nu} = -T^a_{\nu\mu}$

$$\omega'^a{}_\mu{}^b = \Lambda^{-1 a}{}_c \partial_\mu \Lambda^c{}_b + \Lambda^{-1 a}{}_c \omega_\mu{}^c{}_d \Lambda^d{}_b.$$

same transformation properties that YM potential for the group  $O(D-1,1)$

# First structure equation

- Lorentz Covariant derivatives

$$V'^a(x) = \Lambda^{-1 a}{}_b(x)V^b(x),$$

$$U'_a(x) = U_b(x)\Lambda^b{}_a(x),$$

$$T'_{ab}(x) = T_{cd}(x)\Lambda^c{}_a(x)\Lambda^d{}_b(x),$$

$$D_\mu V^a = \partial_\mu V^a + \omega_\mu{}^a{}_b V^b,$$

$$D_\mu U_a = \partial_\mu U_a - U_b \omega_\mu{}^b{}_a = \partial_\mu U_a + \omega_{\mu a}{}^b U_b,$$

$$D_\mu T_{ab} = \partial_\mu T_{ab} - T_{cb} \omega_\mu{}^c{}_a - T_{ac} \omega_\mu{}^c{}_b.$$

$$D_\mu \eta_{ab} = -\eta_{cb} \omega_\mu{}^c{}_a - \eta_{ac} \omega_\mu{}^c{}_b = -\omega_{\mu ba} - \omega_{\mu ab} = 0.$$

The metric has vanishing covariant derivative.

# First structure equation

scalar products  $V^a U_a$  are preserved under parallel transport

$$\tilde{V}^a(x) = V^a(x) + \omega_{\mu b}^a V^b(x) \Delta x^\mu$$

The geometrical effect of torsion is seen in the properties of an infinitesimal parallelogram constructed by the parallel transport of two vector fields.

For the Levi-Civita connection the torsion vanishes

$$de^a + \omega^a_b \wedge e^b = 0$$

Non-vanishing torsion appears in supergravity

# First structure equation

- Covariant derivatives

$\omega_{\mu[\rho\nu]} = \omega_{\mu ab} e_{\rho}^a e_{\nu}^b$  in terms of

$$\Omega_{[\mu\nu]\rho} = (\partial_{\mu} e_{\nu}^a - \partial_{\nu} e_{\mu}^a) e_{a\rho}, \quad T_{[\mu\nu]\rho} = T_{\mu\nu}{}^a e_{a\rho}$$

The structure equation  $T_{[\mu\nu]\rho} = \Omega_{[\mu\nu]\rho} + \omega_{\mu[\rho\nu]} - \omega_{\nu[\rho\mu]}$  implies

$$\begin{aligned} \omega_{\mu[\nu\rho]} &= \omega_{\mu[\nu\rho]}(e) + K_{\mu[\nu\rho]}, \\ \omega_{\mu[\nu\rho]}(e) &= \frac{1}{2}(\Omega_{[\mu\nu]\rho} - \Omega_{[\nu\rho]\mu} + \Omega_{[\rho\mu]\nu}) = \omega_{\mu ab}(e) e_{\nu}^a e_{\rho}^b, \\ \omega_{\mu}{}^{ab}(e) &= 2e^{\nu[a} \partial_{[\mu} e_{\nu]}^{b]} - e^{\nu[a} e^{b]\sigma} e_{\mu c} \partial_{\nu} e_{\sigma}{}^c, \\ K_{\mu[\nu\rho]} &= -\frac{1}{2}(T_{[\mu\nu]\rho} - T_{[\nu\rho]\mu} + T_{[\rho\mu]\nu}). \end{aligned}$$

$K_{\mu[\nu\rho]}$  is called contorsion

# First structure equation

- Covariant derivatives

Matrix representation of Lorentz transformation

$$\Lambda(x) = \exp\left(\frac{1}{2}\lambda^{ab}(x)m_{\{ab\}}\right)$$

Spinor representation  $\Psi'(x) = \exp\left(-\frac{1}{4}\lambda^{ab}(x)\gamma_{ab}\right)\Psi(x),$

$$D_\mu\Psi(x) = \left(\partial_\mu + \frac{1}{4}\omega_\mu{}^{ab}(x)\gamma_{ab}\right)\Psi(x).$$

# First structure equation

- The affine connection

Our next task is to transform Lorentz covariant derivatives to covariant derivatives with respect to general conformal transformations

$$\begin{aligned}\nabla_{\mu} V^{\rho} &\equiv e_a^{\rho} D_{\mu} V^a \\ &= e_a^{\rho} D_{\mu} (e_{\nu}^a V^{\nu}) \\ &= \partial_{\mu} V^{\rho} + e_a^{\rho} (\partial_{\mu} e_{\nu}^a + \omega_{\mu}{}^a{}_b e_{\nu}^b) V^{\nu} .\end{aligned}$$

Affine connection  $\Gamma_{\mu\nu}^{\rho} = e_a^{\rho} (\partial_{\mu} e_{\nu}^a + \omega_{\mu}{}^a{}_b e_{\nu}^b) .$

relates affine connection with spin connection

$$\begin{aligned}\nabla_{\mu} V^{\rho} &= \partial_{\mu} V^{\rho} + \Gamma_{\mu\nu}^{\rho} V^{\nu} \\ \partial_{\mu} e_{\nu}^a + \omega_{\mu}{}^a{}_b e_{\nu}^b - \Gamma_{\mu\nu}^{\sigma} e_{\sigma}^a &= 0 \quad \text{vielbein postulate}\end{aligned}$$

# First structure equation

$$\nabla_\mu V_\nu \equiv e_\nu^a D_\mu V_a \quad \nabla_\mu V_\nu = \partial_\mu V_\nu - e_a^\rho (\partial_\mu e_\nu^a + \omega_\mu^a{}_b e_\nu^b) V_\rho.$$

$$\nabla_\mu V_\nu = \partial_\mu V_\nu - \Gamma_{\mu\nu}^\rho V_\rho$$

For tensors in general  $\nabla_\mu T_{\nu_1 \dots \nu_q}^{\rho_1 \dots \rho_p} \equiv e_{a_1}^{\rho_1} \dots e_{a_p}^{\rho_p} e_{\nu_1}^{b_1} \dots e_{\nu_q}^{b_q} D_\mu T_{b_1 \dots b_q}^{a_1 \dots a_p}$

$$\nabla_\mu g_{\nu\rho} \equiv \partial_\mu g_{\nu\rho} - \Gamma_{\mu\nu}^\sigma g_{\sigma\rho} - \Gamma_{\mu\rho}^\sigma g_{\nu\sigma} = 0$$

Covariant differentiation commutes with index raising

$$\nabla_\mu V^\rho = g^{\rho\nu} \nabla_\mu V_\nu$$

# First structure equation

- The affine connection

$$\Gamma_{\mu\nu}^{\rho} = \Gamma_{\mu\nu}^{\rho}(g) - K_{\mu\nu}^{\rho},$$
$$\Gamma_{\mu\nu}^{\rho}(g) = \frac{1}{2}g^{\rho\sigma}(\partial_{\mu}g_{\sigma\nu} + \partial_{\nu}g_{\mu\sigma} - \partial_{\sigma}g_{\mu\nu}).$$

$$\Gamma_{\mu\nu}^{\rho} - \Gamma_{\nu\mu}^{\rho} = -K_{\mu\nu}^{\rho} + K_{\nu\mu}^{\rho} = T_{\mu\nu}^{\rho}.$$

For mixed quantities with both coordinate and frame indexes, it is useful to distinguish among local Lorentz and coordinate covariant derivatives

$$D_{\mu}\Psi_{\nu} \equiv \left(\partial_{\mu} + \frac{1}{4}\omega_{\mu ab}\gamma^{ab}\right)\Psi_{\nu},$$

$$\nabla_{\mu}\Psi_{\nu} = D_{\mu}\Psi_{\nu} - \Gamma_{\mu\nu}^{\rho}\Psi_{\rho}.$$

Vielbein postulate equivalent to  $\nabla_{\mu}e_{\nu}^a = \partial_{\mu}e_{\nu}^a + \omega_{\mu}{}^a{}_b e_{\nu}^b - \Gamma_{\mu\nu}^{\sigma}e_{\sigma}^a = 0$



# First structure equation

- **Partial integration** We have

$$\partial_\mu \sqrt{-g} = \sqrt{-g} \Gamma_{\rho\mu}^\rho(g) \quad \text{from which}$$

$$\int d^D x \sqrt{-g} \nabla_\mu V^\mu = \int d^D x \partial_\mu (\sqrt{-g} V^\mu) - \int d^D x \sqrt{-g} K_{\nu\mu}{}^\nu V^\mu$$

The second term shows the violation of the manipulations of the integration by Parts in the case of torsion

$$K_{\nu\mu}{}^\nu = -T_{\nu\mu}{}^\nu$$

# Second structure equation

- Curvature tensor

spin connection  $\omega_{\mu ab}$  transforms as a YM gauge potential for the Group  $O(D-1,1)$

$$R_{\mu\nu ab} \equiv \partial_\mu \omega_{\nu ab} - \partial_\nu \omega_{\mu ab} + \omega_{\mu ac} \omega_\nu{}^c{}_b - \omega_{\nu ac} \omega_\mu{}^c{}_b$$

YM field strength. We define the curvature two form

$$\rho^{ab} = \frac{1}{2} R_{\mu\nu}{}^{ab}(x) dx^\mu \wedge dx^\nu .$$

Second structure equation

$$d\omega^{ab} + \omega^a{}_c \wedge \omega^{cb} = \rho^{ab} .$$

# Bianchi identities

$$\begin{aligned}\rho^{ab} \wedge e_b &= dT^a + \omega^{ab} \wedge T_b, \\ d\rho^{ab} + \omega^a_c \wedge \rho^{cb} - \rho^{ac} \wedge \omega_c^b &= 0\end{aligned}$$

using  $R_{\mu\nu\rho}^a = R_{\mu\nu b}^a e_\rho^b$  we have

$$R_{\mu\nu\rho}^a + R_{\nu\rho\mu}^a + R_{\rho\mu\nu}^a = -D_\mu T_{\nu\rho}^a - D_\nu T_{\rho\mu}^a - D_\rho T_{\mu\nu}^a$$

First Bianchi identity, it has no analogue in YM

$$D_\mu R_{\nu\rho}^{ab} + D_\nu R_{\rho\mu}^{ab} + D_\rho R_{\mu\nu}^{ab} = 0$$

usual Bianchi identity for YM

useful relation

$$\delta R_{\mu\nu ab} = D_\mu \delta \omega_{\nu ab} - D_\nu \delta \omega_{\mu ab}$$

# Ricci identities and curvature tensor

Commutator of covariant derivatives

$$\begin{aligned}[D_\mu, D_\nu]\Phi &= \frac{1}{2}R_{\mu\nu ab}M^{ab}\Phi \\ [D_\mu, D_\nu]V^a &= R_{\mu\nu}{}^a{}_b V^b, \\ [D_\mu, D_\nu]\Psi &= \frac{1}{4}R_{\mu\nu ab}\gamma^{ab}\Psi.\end{aligned}$$

$$[\nabla_\mu, \nabla_\nu]V^\rho = R_{\mu\nu}{}^\rho{}_\sigma V^\sigma - T_{\mu\nu}{}^\sigma \nabla_\sigma V^\rho$$

Curvature tensor

$$R_{\mu\nu}{}^\rho{}_\sigma = \partial_\mu \Gamma_{\nu\sigma}^\rho - \partial_\nu \Gamma_{\mu\sigma}^\rho + \Gamma_{\mu\tau}^\rho \Gamma_{\nu\sigma}^\tau - \Gamma_{\nu\tau}^\rho \Gamma_{\mu\sigma}^\tau$$

$$R_{\mu\nu}{}^\rho{}_\sigma = R_{\mu\nu ab}e^{a\rho}e_\sigma^b$$

Second Bianchi identity

$$\nabla_\mu R_{\nu\rho}{}^{\sigma\tau} + \nabla_\nu R_{\rho\mu}{}^{\sigma\tau} + \nabla_\rho R_{\mu\nu}{}^{\sigma\tau} = T_{\mu\nu}{}^\xi R_{\xi\rho}{}^{\sigma\tau} + T_{\nu\rho}{}^\xi R_{\xi\mu}{}^{\sigma\tau} + T_{\rho\mu}{}^\xi R_{\xi\nu}{}^{\sigma\tau}$$

# Ricci tensor

Ricci tensor  $R_{\mu\nu} = R_{\mu}{}^{\sigma}{}_{\nu\sigma}$

Scalar curvature  $R = g^{\mu\nu} R_{\mu\nu}$

If there is no torsion  $R_{\mu\nu} = R_{\nu\mu}$   $R_{\mu\nu\rho\sigma} = R_{\rho\sigma\mu\nu}$

Useful relation  $\delta R_{\mu\nu}{}^{\rho}{}_{\sigma} = \nabla_{\mu}\delta\Gamma_{\nu\sigma}^{\rho} - \nabla_{\nu}\delta\Gamma_{\mu\sigma}^{\rho}$

Hilbert action

$$\frac{1}{(D-2)!} \int \varepsilon_{abc_1\dots c_{D-2}} e^{c_1} \wedge \dots \wedge e^{c_{D-2}} \wedge \rho^{ab} = \int d^D x \sqrt{-g} R$$

# Dimensional analysis and Planck units

$$\frac{Gm_p}{c^2} = \frac{\hbar}{m_p c}$$

$$m_p = \sqrt{\frac{\hbar c}{G}} \simeq 10^{-5} \text{ grams} \simeq 10^{19} \text{ GeV}$$

$$l_p = \sqrt{\frac{G\hbar}{c^3}} = \sqrt{\frac{\kappa^2}{8\pi}} \simeq 10^{-33} \text{ cm}$$

$$t_p = \sqrt{\frac{G\hbar}{c^5}} = \sqrt{\frac{\kappa^2}{8\pi c^2}} \simeq 10^{-44} \text{ seconds}$$

# The first and second order formulations of general relativity

- Second order formalism

Field content,  $g_{\mu\nu}(x)$   $\phi(x)$   $A_\mu(x)$

Action

$$S = \int d^D x \sqrt{-\det g} \left( \frac{1}{2\kappa^2} g^{\mu\nu} R_{\mu\nu}(g) + L \right)$$
$$L = -\frac{1}{2} g^{\mu\nu} \partial_\mu \phi \partial_\nu \phi - \frac{1}{4} g^{\mu\rho} g^{\nu\sigma} F_{\mu\nu} F_{\rho\sigma} .$$

$\kappa^2 = 8\pi G_N$  is the gravitational coupling constant

$$M_P \equiv M_{\text{Planck}} / \sqrt{8\pi} \sim 2 \times 10^{18} \text{ GeV}$$

$$\partial_\mu \phi = \nabla_\mu \phi \quad F_{\mu\nu} = \nabla_\mu A_\nu - \nabla_\nu A_\mu = \partial_\mu A_\nu(x) - \partial_\nu A_\mu(x)$$

# The first and second order formulations of general relativity

- Variation of the action

$$\delta S = \frac{1}{2\kappa^2} \int d^D x [\delta(\sqrt{-\det g} g^{\mu\nu}) R_{\mu\nu} + \sqrt{-\det g} g^{\mu\nu} \delta R_{\mu\nu} + \dots]$$

Last term total derivative due  $\delta R_{\mu\nu} = \nabla_\rho \delta \Gamma_{\mu\nu}^\rho - \nabla_\mu \delta \Gamma_{\nu\rho}^\rho$  plus no torsion

Einstein equations

$$G_{\mu\nu} \equiv R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = \kappa^2 T_{\mu\nu},$$

$$T_{\mu\nu} \equiv -2 \frac{1}{\sqrt{-\det g}} \frac{\delta(\sqrt{-\det g} L)}{\delta g^{\mu\nu}} = \partial_\mu \phi \partial_\nu \phi + F_\mu{}^\rho F_{\nu\rho} + g_{\mu\nu} L$$

The Einstein equations are consistent only if the matter tensor  $\nabla^\mu T_{\mu\nu} = 0$



# The first and second order formulations of general relativity

- Conservation energy-momentum tensor

The invariance under diff of the matter action

$$\delta \int d^D x \sqrt{-\det g} L = \int d^D x \sqrt{-\det g} \left[ \frac{1}{2} T^{\mu\nu} (\nabla_\mu \xi_\nu + \nabla_\nu \xi_\mu) + \frac{\delta L}{\delta \phi} \mathcal{L}_\xi \phi + \frac{\delta L}{\delta A_\mu} \mathcal{L}_\xi A_\mu \right] = 0$$

if  $\phi$  and  $A_\mu$  do satisfy the equations of motion

$$\int d^D x \sqrt{-\det g} T^{\mu\nu} \nabla_\mu \xi_\nu = 0$$

which implies

$$\nabla^\mu T_{\mu\nu} = 0$$

# The first and second order formulations of general relativity

- Scalar and gauge field equations

$$\begin{aligned}\partial_\mu(\sqrt{-\det g} g^{\mu\nu} \partial_\nu \phi) &= \sqrt{-\det g} g^{\mu\nu} \nabla_\mu \partial_\nu \phi = 0 \\ \partial_\mu(\sqrt{-\det g} F^{\mu\nu}) &= \sqrt{-\det g} \nabla_\mu F^{\mu\nu} = 0.\end{aligned}$$

Ricci form of Einstein field equation

$$R_{\mu\nu} = \kappa^2 \left[ T_{\mu\nu} - \frac{1}{D-2} g_{\mu\nu} T^\rho{}_\rho \right]$$

# The first and second order formulations of general relativity

## Useful relations

$$\begin{aligned}\delta g^{\mu\nu} &= -g^{\mu\rho} \delta g_{\rho\sigma} g^{\sigma\nu}, \\ \delta \sqrt{-\det g} &= \frac{1}{2} \sqrt{-\det g} g^{\mu\nu} \delta g_{\mu\nu} = -\frac{1}{2} \sqrt{-\det g} g_{\mu\nu} \delta g^{\mu\nu}\end{aligned}$$

Particular case of

$$\delta \det M = (\det M) \operatorname{Tr}(M^{-1} \delta M)$$

# The first and second order formulations of general relativity

- Matter scalars

$$\mathcal{L} = -\frac{1}{2}\partial_\mu\phi\partial^\mu\phi - V(\phi)$$

$$R_{\mu\nu} = \kappa^2 \left[ \partial_\mu\phi\partial_\nu\phi + g_{\mu\nu}\frac{2}{D-2}V \right]$$

# The first and second order formulations of general relativity

## Gravitational fluctuations of flat spacetime

In absence of matter  $R_{\mu\nu} = 0$  Solution  $g_{\mu\nu} = \eta_{\mu\nu}$

fluctuations  $g_{\mu\nu}(x) = \eta_{\mu\nu} + \kappa h_{\mu\nu}(x)$

$$\Gamma_{\mu\nu}^{\rho \text{Lin}} = \frac{1}{2} \kappa \eta^{\rho\sigma} (\partial_{\mu} h_{\sigma\nu} + \partial_{\nu} h_{\mu\sigma} - \partial_{\sigma} h_{\mu\nu})$$

$$R_{\mu\nu}^{\text{Lin}} = -\frac{1}{2} \kappa [\square h_{\mu\nu} - \partial^{\rho} (\partial_{\mu} h_{\rho\nu} + \partial_{\nu} h_{\mu\rho}) + \partial_{\mu} \partial_{\nu} h^{\rho}_{\rho}]$$

The gauge transformations are obtained linearizing the diff transformations

$$\begin{aligned} \delta g_{\mu\nu} &= \kappa (\nabla_{\mu} \xi_{\nu} + \nabla_{\nu} \xi_{\mu}), \\ \delta R_{\mu\nu} &= \kappa (\xi^{\rho} \nabla_{\rho} R_{\mu\nu} + \nabla_{\mu} \xi^{\rho} R_{\rho\nu} + \nabla_{\nu} \xi^{\rho} R_{\mu\rho}) \end{aligned}$$

# The first and second order formulations of general relativity

$$\begin{aligned}\delta h_{\mu\nu} &= \partial_\mu \xi_\nu + \partial_\nu \xi_\mu \\ \delta R_{\mu\nu}^{\text{Lin}} &= 0.\end{aligned}$$

Degrees of freedom. Choose the gauge  $\partial^i h_{i\mu} = 0$

Fixes completely the gauge from

$$\delta(\partial^i h_{i\mu}) = \nabla^2 \xi_\mu + \partial^i \partial_\mu \xi_i = 0$$

Let  $\mu \rightarrow j$  and contract with  $\partial^j$  to learn that  $\nabla^2 \partial^i \xi_i = 0$

We have  $\xi_\mu(x) \equiv 0$

The equations of motion become

$$\square h_{\mu\nu} - \partial^0(\partial_\mu h_{0\nu} + \partial_\nu h_{\mu 0}) + \partial_\mu \partial_\nu (h_{ii} - h_{00}) = 0.$$

# The first and second order formulations of general relativity

- Degree of freedom

$$\mu = \nu = 0 \quad : \quad \nabla^2 h_{00} + \partial_0^2 h_{ii} = 0,$$

$$\mu = \nu = i \quad : \quad 2\nabla^2 h_{ii} - \partial_0^2 h_{ii} - \nabla^2 h_{00} = 0$$

$$\nabla^2 h_{ii} = 0, \text{ so that } h_{ii} = 0 \quad h_{00} \equiv 0 \quad h_{0i} \equiv 0 \quad \text{oi term}$$

The non-trivial equations are  $\square h_{ij} = 0$

Constraints  $\partial^i h_{ij} = 0 \quad h_{ii} = 0.$

Since  $D(D-1)/2 - (D-1) - 1 = D(D-3)/2$

The number of on-shell degrees of freedom, helicities is  $D(D-3)/2$

$\square$  Symmetric traceless representation

# The first and second order formulations of general relativity

## Second order formalism for gravity and fermions

Field content, frame field  $e_\mu^a(x)$  massless Dirac field  $\Psi(x)$

Spin connection dependent quantity  $\omega_\mu^{ab}(e)$

$$S = S_2 + S_{1/2} = \int d^D x e \left[ \frac{1}{2\kappa^2} e_\mu^a e_\nu^b R_{\mu\nu}{}^{ab}(e) - \frac{1}{2} \bar{\Psi} \gamma^\mu \nabla_\mu \Psi + \frac{1}{2} \bar{\Psi} \overleftarrow{\nabla}_\mu \gamma^\mu \Psi \right]$$

$$R_{\mu\nu ab} = \partial_\mu \omega_{\nu ab} - \partial_\nu \omega_{\mu ab} + \omega_{\mu ac} \omega_\nu{}^c{}_b - \omega_{\nu ac} \omega_\mu{}^c{}_b$$

$$S_2(e, \omega) \quad \left. \frac{\delta S_2}{\delta \omega_{\mu ab}} \right|_{\omega=\omega(e)} = 0$$



# The first and second order formulations of general relativity

$$\begin{aligned}\nabla_\mu \Psi &= D_\mu \Psi = \left( \partial_\mu + \frac{1}{4} \omega_\mu^{ab} \gamma_{ab} \right) \Psi \\ \bar{\Psi} \overleftarrow{\nabla}_\mu &= \bar{\Psi} \overleftarrow{D}_\mu = \bar{\Psi} \left( \overleftarrow{\partial}_\mu - \frac{1}{4} \omega_\mu^{ab} \gamma_{ab} \right)\end{aligned}$$

The total covariant derivative and the Lorentz covariant derivative coincide for spinor field but not for the gravitino

# The first and second order formulations of general relativity

- Curved space gamma matrices

Constant gamma matrices verify  $\{\gamma^a, \gamma^b\} = 2\eta^{ab}$

Frame fields are used to transform frame *vector* indices to a coordinate basis.

$$\gamma^\mu = e_a^\mu \gamma^a = g^{\mu\nu} \gamma_\nu \quad \{\gamma^\mu, \gamma^\nu\} = g^{\mu\nu}$$

The curved gamma matrices transform a vector under coordinate transformations  
But they have also spinor indexes

$$\begin{aligned} \nabla_\mu \gamma_\nu &= \partial_\mu \gamma_\nu + \frac{1}{4} \omega_\mu^{ab} [\gamma_{ab}, \gamma_\nu] - \Gamma_{\mu\nu}^\rho \gamma_\rho \\ &= \gamma^a (\partial_\mu e_{a\nu} + \omega_{\mu ab} e_\nu^b - \Gamma_{\mu\nu}^\rho e_{a\rho}) \end{aligned}$$

$$\nabla_\mu \gamma_\nu = 0$$

holds for any affine connection with or without torsion

# The first and second order formulations of general relativity

- Fermion equation of motion

$$\gamma^\mu \nabla_\mu \Psi = 0$$

$$\gamma^\mu \nabla_\mu \gamma^\nu \nabla_\nu \Psi = \left( g^{\mu\nu} \nabla_\mu \nabla_\nu - \frac{1}{4} R \right) \Psi$$

Steps in the derivation of Einstein equation

$$\delta S = \int d^D x e \left[ \frac{1}{\kappa^2} \left( e^{b\nu} R_{\mu\nu ab} - \frac{1}{2} e_{a\mu} R \right) \delta e^{a\mu} - \frac{1}{2} \bar{\Psi} \gamma^a \overleftrightarrow{\nabla}_\mu \Psi \delta e^{a\mu} - \frac{1}{8} \bar{\Psi} \{ \gamma^\mu, \gamma^{ab} \} \Psi \delta \omega_{\mu ab} \right]$$

We drop a term proportional to the fermion lagrangian because we use the equations of motion for the fermion

# The first and second order formulations of general relativity

$$\delta S = \int d^D x \left[ \frac{1}{\kappa^2} \left( e^{b\nu} R_{\mu\nu ab} - \frac{1}{2} e_{a\mu} R \right) \delta e^{a\mu} - \frac{1}{4} \bar{\Psi} \left[ \gamma_a \overleftrightarrow{\nabla}_\mu + \gamma_\mu e_a^\rho \overleftrightarrow{\nabla}_\rho \right] \Psi \delta e^{a\mu} \right]$$

From which we deduce the Einstein equation

$$R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = \kappa^2 T_{\mu\nu} \equiv -\kappa^2 \frac{1}{4} \bar{\Psi} \left[ \gamma_\mu \overleftrightarrow{\nabla}_\nu + \gamma_\nu \overleftrightarrow{\nabla}_\mu \right] \Psi$$

The stress tensor is the covariant version of flat Dirac symmetric stress tensor

$$\nabla^\mu T_{\mu\nu} = 0, \text{ and symmetry, } T_{\mu\nu} = T_{\nu\mu}$$

Follows from the matter being invariant coordinate and local Lorentz transformations

# The first and second order formulations of general relativity

- The first order formalism for gravity and fermions

Field content    frame field  $e_\mu^a$  and spin connection  $\omega_{\mu ab}$

Fermion field     $\Psi(x)$     Same form of the action as in the second order formalism but now vielbein and spin conn independent

Variation of the gravitational action

$$\delta S_2 = \frac{1}{2\kappa^2} \int d^D x e e_\mu^a e_\nu^b \left( D_\mu \delta \omega_\nu^{ab} - D_\nu \delta \omega_\mu^{ab} \right)$$

We have used     $\delta R_{\mu\nu ab} = D_\mu \delta \omega_\nu^{ab} - D_\nu \delta \omega_\mu^{ab}$

$$\delta S_2 = \frac{1}{2\kappa^2} \int d^D x e e_\mu^a e_\nu^b \left( 2\nabla_\mu \delta \omega_\nu^{ab} + T_{\mu\nu}^\rho \delta \omega_\rho^{ab} \right)$$

# The first and second order formulations of general relativity

- The first order formalism for gravity and fermions

Integration by parts

$$\int d^D x \sqrt{-g} \nabla_\mu V^\mu = \int d^D x \partial_\mu (\sqrt{-g} V^\mu) - \int d^D x \sqrt{-g} K_{\nu\mu}{}^\nu V^\mu$$

$$\delta S_2 = \frac{1}{2\kappa^2} \int d^D x e \left( -2K_{\rho\mu}{}^\rho e_a^\mu e_b^\nu \delta\omega_\nu{}^{ab} + T_{ab}{}^\rho \delta\omega_\rho{}^{ab} \right)$$

$$= \frac{1}{2\kappa^2} \int d^D x e \left( T_{\rho a}{}^\rho e_b^\nu - T_{\rho b}{}^\rho e_a^\nu + T_{ab}{}^\nu \right) \delta\omega_\nu{}^{ab},$$

Form the fermion action

$$\delta S_{1/2} = -\frac{1}{8} \int d^D x e \bar{\Psi} \{ \gamma^\nu, \gamma_{ab} \} \Psi \delta\omega_\nu{}^{ab}$$

$$= -\frac{1}{4} \int d^D x e \bar{\Psi} \gamma^\nu{}_{ab} \Psi \delta\omega_\nu{}^{ab}.$$

# The first and second order formulations of general relativity

- The first order formalism for gravity and fermions

The equations of motion of the spin connection gives

$$T_{ab}{}^\nu - T_{a\rho}{}^\rho e_b^\nu + T_{b\rho}{}^\rho e_a^\nu = \frac{1}{2}\kappa^2 \bar{\Psi} \gamma_{ab}{}^\nu \Psi$$

the right hand side is traceless therefore also the torsion is traceless

$$T_{ab}{}^\nu = \frac{1}{2}\kappa^2 \bar{\Psi} \gamma_{ab}{}^\nu \Psi = -2K^\nu{}_{ab}$$

If we substitute  $\omega = \omega(e) + K$

$$S = \frac{1}{2\kappa^2} \int d^D x e \left[ R(g) - \kappa^2 \bar{\Psi} \gamma^\mu \overleftrightarrow{\nabla}_\mu \Psi \right. \quad (1) \\ \left. - 2\nabla_\mu K_\nu{}^{\nu\mu} + K_{\mu\nu\rho} K^{\nu\mu\rho} - K_\rho{}^\rho{}_\mu K_\sigma{}^{\sigma\mu} - \frac{1}{2} \bar{\Psi} \gamma_{\mu\nu\rho} \Psi K^{\mu\nu\rho} \right]$$

# The first and second order formulations of general relativity

- The first order formalism for gravity and fermions

The physical equivalent second order action is

$$S = \frac{1}{2} \int d^D x e \left[ \frac{1}{\kappa^2} R(g) - \bar{\Psi} \gamma^\mu \overleftrightarrow{\nabla}_\mu \Psi + \frac{1}{16} \kappa^2 (\bar{\Psi} \gamma_{\mu\nu\rho} \Psi) (\bar{\Psi} \gamma^{\mu\nu\rho} \Psi) \right]$$

Physical effects in the fermion theories with torsion and without torsion

Differ only in the presence of quartic fermion term.

This term generates 4-point contact diagrams .